SuperKEKB

February 8, 2007

KEK Topical Conference

Masa Yamauchi KEK

Outline

Introduction

- Achievements of the B factories and the next step
- Physics at SuperKEKB
- Two possible approaches to achieve $L \sim 10^{36}$
 - High current approach: SuperKEKB
 - Low β and low ϵ approach: LC inspired SuperB
- Schedule and funding status
- Detector issues
- Summary





 e^{-} B factory with much higher L_{peak} .

What is next with **B** physics?

- If new physics at O(1)TeV...
 - It is natural to assume that the effects are seen in $B/D/\tau$ decays.
 - Flavour structure of new physics?
 - CP violation in new physics?
 - These studies will be useful to identify mechanism of SUSY breaking, if NP=SUSY.

Otherwise...

 Search for deviations from SM in flavor physics will be one of the best ways to find new physics.



Likelihood for the effects of new physics to be seen in *B* decays.

SuperKEKB

- Asymmetric energy e⁺e⁻ collider at E_{CM}=m(Y(4S)) to be realized by upgrading the existing KEKB collider.
- Super-high luminosity $\approx 8 \times 10^{35}$ /cm²/sec $\rightarrow 1 \times 10^{10}$ BB per yr.

 \rightarrow 8×10 ⁹ τ + τ - per yr.







Search for new CP phases

In general, new physics contains new sources of flavor mixing and CP violation.

▶ In SUSY models, for example, SUSY particles contribute to the $b \rightarrow s$ transition, and their CP phases change CPV observed in $B \rightarrow \phi K$, $\eta' K$ etc.



Sensitivity to new CP phases Estimated error in the measurement of Discovery region with 50 ab-1 time dependent CP violation 1 3σ Discovery Region 0.0 0.5 0.4 0.8 0.3 0.2 |A_{NP}/A_{SM}|² 9.0 Error on AS 0.06 0.05 0.2 0.04 0.03 0 0.02 -3 -2 -1 0 10.1 10 $\theta_{\text{new physics}}$ Integrated luminosity (ab⁻¹)

A possible hint for NP: $b \rightarrow s \overline{q} q$



Search for new flavor mixing



Possible observables:

: Probe the flavor changing process with the "EW probe".

This measurement is especially sensitive to new physics such as SUSY, heavy Higgs and extra dim.

Theoretical predictions for l^+l^- forward-backward charge asymmetry for SM and SUSY model with various parameter sets.



--SM -- CLL=3 -- C7=-C7SM

The F/B asymmetry is a consequence of γ - Z^0 interference.

Sensitivity to new flavor mixing



"B meson beam" technique







Search for flavor-violating τ decay





Comparison with LHCb

<i>e</i> + <i>e</i> - is advantageous in…	LHCb is advantageous in	
CPV in $B \rightarrow \phi K_S$, $\eta' K_S$, CPV in $B \rightarrow K_S \pi^0 \gamma$ $B \rightarrow K \nu \nu$, $\tau \nu$, $D^{(*)} \tau \nu$	CPV in $B \rightarrow J/\psi K_S$ Most of <i>B</i> decays not including ν or γ	
Inclusive $b \rightarrow s \mu \mu$, see $\tau \rightarrow \mu \gamma$ and other LFV $D^0 \overline{D^0}$ mixing	Time dependent measurements of B_S $B_{(S,d)} \rightarrow \mu\mu$ B_{a} and bottomed baryons	

These are complementary to each other !!



High current approach – Technical challenge

- Powerful RF system needed to store high current – Larger wall-plug power
- High beam-beam parameter with head-on collision
- Highly nonlinear system
- Sensitive for noise and error due to high beambeam parameter
- Limit of bunch length, instability and coherent synchrotron radiation



Parameter sets for SuperB

	KEKB	SuperKEKB	LC-SuperB
εx	18(L)/24(H) ×10 ⁻⁹	9.0×10 ⁻⁹	0.8×10 ⁻⁹
εγ	18(L)/24(H) ×10 ⁻¹¹	4.5×10 ⁻¹¹	2×10 ⁻¹²
βx (mm)	560	200	9
βy (mm)	6	3	0.133
σz (mm)	7	3	6
νs	0.025	0.025	0.012/0.026
ne	(1388×) 5.5×10 ¹⁰	(5000×) 5.5×10 ¹⁰	(5000×) 1.9×10 ¹⁰
np	(1388×) 7.5×10 ¹⁰	(5000×) 1.26×10 ¹¹	(5000×) 3.3×10 ¹⁰
θ /2 (mrad)	11	0 (crab crossing)	25
ξx	0.1	0.397	
ξy	0.06	0.794->0.24	
Lum.	1.7×10 ³⁴	8×10 ³⁵	1×10 ³⁶

Crab cavity





(Strong-strong simulation)

crossing angle 30 mrad (at the optimum tune)

•Superconducting crab cavities are under development, and will be installed in KEKB in 2006.

K. Ohmi





Oct. 16, 2006



A pair of crab cavities will be installed in KEKB in Jan. 07.

KEKB's approach

We keep high current approach until finding a solution to dynamic aperture and injection scheme in the low emittance scheme.

Both crab crossing and crab waist scheme will be tested in the KEKB ring soon.





The beam pipes and all vacuum components will be replaced with higher-current-proof design.

will reach 8 × 10^{35} cm⁻²s⁻¹.

Proposed schedule





However INFN is a strong supporter in this enterprise.

INFN will try to gain the support of italian government for such an international project.

Conclusion 2

Only one Super B Factory will be built if one. Site decision should be open Lumi as high as possible (> 10 ³⁶ better 10 ⁴³ !!) We have set up an International Group Collaboration between SuperB and SuperKEKB is needed and welcome!

Requirements to the detector

Issues

- Higher background (×20)
 - radiation damage and occupancy
 - fake hits and pile-up noise in the EM
- Higher event rate (×50)
 - higher rate trigger, DAQ and computing
- Require special features
 - low $p \mu$ identification \leftarrow s $\mu\mu$ recon. eff.
 - hermeticity $\leftarrow v$ "reconstruction"

Possible solution:

- Replace inner layers of the vertex detector with a silicon striplet detector.
- Replace inner part of the central tracker with a silicon strip detector.
- Better particle identification device
- ▶ Replace endcap calorimeter by pure Csl.
- Faster readout electronics and computing system.



Super Belle







Summary

- KEKB/Belle and PEP-II/BaBar have been running very successfully, and brought important scientific and technical achievements.
- Next generation e⁺e⁻ B factory with L~10³⁶ will be very useful to study the new sources of flavor mixing and CP violation.
 - Search for new CPV in $b \rightarrow s$ transition
 - Very precise test of CKM scheme
 - Search for lepton flavor violating τ decays
 - Studies of H^{\pm} interactions with fermions
 - Very precise measurements of α_{s} (@10GeV), $\sin^{2}\theta_{W}$ (@10GeV)...
- Two possible approaches:
 - High current + crab crossing → SuperKEKB
 - Low β and low ϵ + crab waist \rightarrow LC SuperB
- SuperKEKB upgrade and LC inspired SuperB have been proposed.

Goal: $\int L dt = 50-100 \text{ ab}^{-1}$

Most of the interesting measurements will be limited by unavoidable systematics when we reach 50-100 ab⁻¹.

Obs.	$\delta_{_{stat}}$ with 50ab ⁻¹	δ_{syst} with 50ab ⁻¹	Theory err.
$sin2\phi_1$	0.004	0.014	~0.01
ϕ_2	1.2°	a few °	
ϕ_3	1.2°	O(1) °	
$ V_{ub} $	1%	~1%	~5 %
$S_{{oldsymbol{\phi}} K {f s}}$	0.023	0.020	
$A_{\phi K \mathrm{s}}$	0.016	0.018	
$S_{\eta'Ks}$	0.013	0.020	
$A_{\eta'Ks}$	0.009	0.017	
DCPV in $b \rightarrow s\gamma$	0.003	0.002	0.003

High current approach

$$\frac{\sqrt{\varepsilon_x \beta_x}}{\theta \sigma_z} > 1 \text{ or } \theta = 0$$



Keep
$$\varepsilon_x$$
, β_x and $\sqrt{\frac{\beta_y}{\varepsilon_y}}$.
 $\varepsilon_y \beta_y \to 0$
 $L \to \infty$

 $\beta_y > \sigma_z$ limits luminosity

High current, Small coupling

Choice of operating point

$$v_x \to +0.5 \quad \xi_y \to \infty \qquad N \to \infty$$

 $L \to \infty$

Low emittance approach

$$\frac{\sqrt{\varepsilon_x \beta_x}}{\theta \sigma_z} < 1$$

$$L \sim \frac{N^2}{\theta \sigma_z \sqrt{\varepsilon_y \beta_y}}$$

$$\xi_x \sim \frac{N}{\theta \sigma_z} \sqrt{\frac{\beta_x}{\varepsilon_x}}$$

$$\xi_y \sim \frac{N}{\theta \sigma_z} \sqrt{\frac{\beta_y}{\varepsilon_y}}$$

$$\beta_y > \frac{\sqrt{\varepsilon_x \beta_x}}{\theta}$$

Keep
$$\sqrt{\frac{\beta_y}{\varepsilon_y}}$$
, $\sqrt{\frac{\beta_x}{\varepsilon_x}}$ and $\frac{\sqrt{\varepsilon_x \beta_x}}{\beta_y}$

$$\varepsilon_{y}\beta_{y} \rightarrow 0$$

$$T \rightarrow \infty$$

- Bunch length is free.
- Small beta and small emittance are required.

Crab crossing and crab waist

- Head-on collision is effectively realized by crab cavity, while crab waist control nonlinear interaction induced by crossing angle.
- For high current approach, either scheme will work(, but not both at once).
- For low emittance approach, only crab waist scheme is applicable.

Beam-beam parameter becomes too high

Consensus in J-HEP community		
First priority is to realize ILC, and its R&D should be boosted. On the other hand, flavor physics programs (SuperKEKB and v at J-PARC) should be carried out as physics program before ILC (i.e., 2010's).		

Ongoing Work



Biagini, Koop, Novohatski, Ohmi, Seeman, Seryi, Shatilov, Sullivan, Raimondi, Wienands, Wolsky, Zobov etc...



Vertical Plane



More

- Low background. Lower than in Babar
- Wall power less than 50 MW(goal 25MW)
- Low beam energy spread for best background rejection as inPEPII or KEKB.
- Upgradable to higher performance
- Possibiliy of reusing existing machine components
- High level of Synergy with other machine projects as ILC.
- " To be studied the possibility to run down to Φ "



Upgradeable



M. Zobov, D. Shatilov

F.Forti - Status of SuperB

A Roodman BNM2006, Sep.06

INFN-AE-05-08 SLAC-PUB-11676

physics/0512235

INFN Process

Workshops Nov 2005 Frascati March 2006 Frascati June 2006 SLAC November 13-15 Rome INFN Roadmap Report SuperB: a linear high-luminosity B Factory Documents INFN Roadmap Report INFN Roadmap Dec 2005 SuperB: a High Luminosity Flavour Factory INFN Roadmap & CERN Strategy Mar 2006 INFN Roadmap SuperB identified as research direction next step is a CDR-type document ~ at end of 2006 see Web site http://www.pi.infn.it/SuperB



Detector issue: background







"Crab waist" (P. Raimondi)



ITEP Meeting 07.25,2006 Moscow

Marcello A Giorgi



2. Bellows Chamber and Gate Valve



Application of Ver.2 to antechamber-type bellows

- Manufactured at BINP (2005)
- Copper cooling channel
 Improve cooling of teeth
- Two bellows chamber were installed into LER wiggler (2005).
- No problem was found up to 1.7 A.







2006 KEKB Review

Three factors to determine luminosity:

